Jake Stowe - Pset I

2.4)

g1 = 9.8;

```
<< PhysicalConstants`
   General::obspkg: PhysicalConstants` is now obsolete. The legacy version being
        loaded may conflict with current functionality. See the Compatibility Guide for updating information.
   2.1)
   a)
   (6.022 * 10 ^ 23) * (10 ^ - 24)
   0.6022
   b)
   WolframAlpha["seconds in a nanocentury", "WolframResult"]
   3.156 s
   (60 * 60 * 24 * 365 * 100 * 10^-9) // N
   3.1536
2.2)
   floppys = (10^12) / (10^6)
   1000000
   Quantity[floppys * 3 / 10 ^ 3, "m"]
   3000 m
2.3)
   Sum[2^i, {i, 0, n}]
   -1 + 2^{1+n}
  2^(1+10^70)
   General::ovfl : Overflow occurred in computation. >>>
```

$$g2 = (6.673 * 10^{-11})$$

6.673 × 10⁻¹¹
decibles = 20 Log[g1/g2]
514.255

2.5)

a)

ChemicalData["TNT"]

ChemicalData["TNT", "MolecularMass"]

227.131 u

$$(1.602*10^-19)$$
 (3) $(6.022*10^23)$ (1000) / (227) 1.27497×10^6

b)

ElementData["Uranium", "KnownIsotopes"]

```
{217, 218, 219, 220, 221, 222, 223, 224, 225, 226, 227, 228, 229, 230, 231, 232, 233, 234, 235, 236, 237, 238, 239, 240, 241, 242}
```

Assume seven divisions 2^7

Solve[10^14 = (1.602 * 10^-19) (10^6) (2^7) (6.022 * 10^23) (1000)
$$\mathbf{x}$$
/(235), \mathbf{x}] {{x \times 1.90307}}

c)

Solve [Quantity [(10^14), "J"] == m Quantity [1, "SpeedofLight"]^2, m] // N
$$\{ m \rightarrow 0.00111265 \text{ kg} \}$$

2.6)

```
a)
velo = Quantity[50, "mph"];
mass = Quantity[100, "g"];
wvlength = UnitConvert[Quantity[1, "PlanckConstant"] / (velo * mass), "meter"]
2.964419 \times 10^{-34} \text{ m}
b)
ChemicalData["Nitrogen"]
N == N
nmass = ChemicalData["Nitrogen", "MolecularMass"]
28.0134 u
tenergy = UnitConvert
  (3/2) (Quantity[1, "BoltzmannConstant"] * Quantity[300, "Kelvins"]), "Joules"]
6.2129 \times 10^{-21} J
nvelo = v / . Solve[(1/2) nmass v^2 = tenergy, v][[2]]
516.84 m/s
nwavelength =
 UnitConvert[Quantity[1, "PlanckConstant"] / (nmass * nvelo), "Angstroms"]
0.27560 Å
c)
FormulaData[{"IdealGasLaw", "Density"}]
mP = (R) T \rho
ndens = \rho / .
  First@Solve[ChemicalData["Nitrogen", "MolarMass"] * Quantity[1, "Atmospheres"] ==
      Quantity[1, "MolarGasConstant"] * Quantity[300, "Kelvins"] * \rho, \rho]
1.13796 \, \text{kg/m}^3
```

ndensconv = UnitConvert[ndens / nmass, "Nanometers" ^ 3]

 $0.024463/nm^3$

nvol = 1 / ndensconv

40.878 nm³

UnitConvert[2*r/. Last@Solve[(4/3)* π *r^3 = nvol, r], "Angstroms"] 42.740 Å

d)

If we assume that the perfect gas law holds at the temperatures we're exploring then we can write UnitConvert[Quantity["MolarGasConstant"], "

$$r[T_{-}] := (1/2) \text{ CubeRoot}[((3/4\pi) * \text{Quantity}[1, "MolarGasConstant"] * Quantity}[1/(6.022 * 10^23), "Moles"] * T)/(Quantity[1, "Atmospheres"])]$$
 Solve[r[t] == a, t]

Solve::ratnz: Solve was unable to solve the system with inexact coefficients.

The answer was obtained by solving a corresponding exact system and numericizing the result. >>

$$\left\{ \left\{ t \rightarrow \text{ConditionalExpression} \left[\left(2.49174 \times 10^{28} \text{ K/m}^3 \right) \text{ Re[a]}^3, \left(\text{Im[a]} = 0 \&\& \text{Re[a]} > 0 \right) \mid \mid \left(\text{Im[a]} = 0 \&\& \text{Re[a]} < 0 \right) \right] \right\}$$

2.4917361702501006`*^28K/m3 * nwavelength^3

0.000521626K

2.7)

a)

Clear[r]

$$Solve[(1/2) m1 v^2 = ((Quantity["GravitationalConstant"] m2 m1)/r), v][[2]]$$

Solve::ratnz: Solve was unable to solve the system with inexact coefficients.

The answer was obtained by solving a corresponding exact system and numericizing the result. \gg

$$\left\{ v \rightarrow \frac{\left(0.00001155 \, \text{m/s} \right) \sqrt{\text{m2} \left(1 \, \text{m}^2 / \left(\text{kg s} \right) \right)} \sqrt{\text{r} \left(1 \, \text{s/m} \right)}}{\text{r}} \right\}$$

Clear[r]

mradius =

```
First@First@(Solve[(1/2) m1 v^2 = ((Quantity["GravitationalConstant"] m2 m1)/r)/. v \rightarrow Quantity["SpeedOfLight"], r]
```

Solve::ratnz: Solve was unable to solve the system with inexact coefficients.

The answer was obtained by solving a corresponding exact system and numericizing the result. >>

$$r \to m2 \left(1.485 \times 10^{-27} \text{ m/kg} \right)$$

c)

mwavelength =

```
First@First@Solve[Quantity["PlanckConstant"] (Quantity["SpeedOfLight"] / \lambda) == m2 Quantity["SpeedOfLight"] ^2, \lambda]
```

Solve::ratnz: Solve was unable to solve the system with inexact coefficients.

The answer was obtained by solving a corresponding exact system and numericizing the result. >>

$$\lambda \to \frac{2.210219 \times 10^{-42} \text{ s}^2\text{N}}{\text{m}^2}$$

Solve::ratnz: Solve was unable to solve the system with inexact coefficients.

The answer was obtained by solving a corresponding exact system and numericizing the result. >>

$$\{m2 \rightarrow 3.858 \times 10^{-8} \text{ kg}\}$$

d)

mradius /. mmass

$$r \rightarrow 5.73 \times 10^{-35} \,\mathrm{m}$$

UnitConvert[Quantity["PlanckLength"], "Meters"]

```
1.616 \times 10^{-35} \, \mathrm{m}
```

f) out of order

mperiod = Quantity["SpeedOfLight"] / (mwavelength[[2]] /. mmass) // UnitSimplify
$$5.23 \times 10^{42}$$
 per second

e)

UnitConvert[Quantity["PlanckConstant"] * mperiod, "Joules"]

```
3.467 \times 10^9 J
```

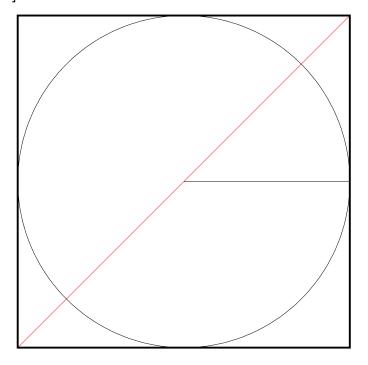
2.8)

a)

By symmetry, the problem is equivalent of the two dimensional case of a triangle and circle

```
{\tt Graphics}\big[\big\{{\tt Sequence}[\,]
```

```
{Line[{{0,0}, {1/2,0}}]
, Red, Line[{{-1/2,-1/2}, {1/2,1/2}}]
}
, {EdgeForm[{Thick, Black}], Circle[{0,0}, 1/2]}
, {Sequence[]
, FaceForm[{Opacity[0]}]
, EdgeForm[{Thick, Black}]
, Rotate[Rectangle[{-1/2,-1/2}], 90 Degree]
}
}
```



Again we see that the 2d Problem is symmetric by rotation of 60 degrees. hence every side is the same length. So we have

$$sol = First@Solve[h = Sqrt[(1/2)^2 + (1/2)^2], h]$$

$$\left\{h \rightarrow \frac{\sqrt{1^2}}{\sqrt{2}}\right\}$$

b)

And we need r of the circle to complete the problem, by inspection we see that

$$rsol = r \rightarrow 1/2$$
$$r \rightarrow \frac{1}{2}$$

pyrvol = Volume[Pyramid[pts28], Assumptions $\rightarrow 1 > 1$]

$$\frac{1^3}{3\sqrt{2}}$$

sphvol =
$$(1/2) * Volume[Ball[{0, 0, 0}, r/. rsol]]$$

 $\frac{1^3 \pi}{12}$

sphvol / pyrvol

$$\frac{\pi}{2\sqrt{2}}$$

 $\label{eq:Graphics3D[{Pyramid[pts28], Ball[{0,0,0}, r/. rsol]} /. 1 \rightarrow 1]}$

